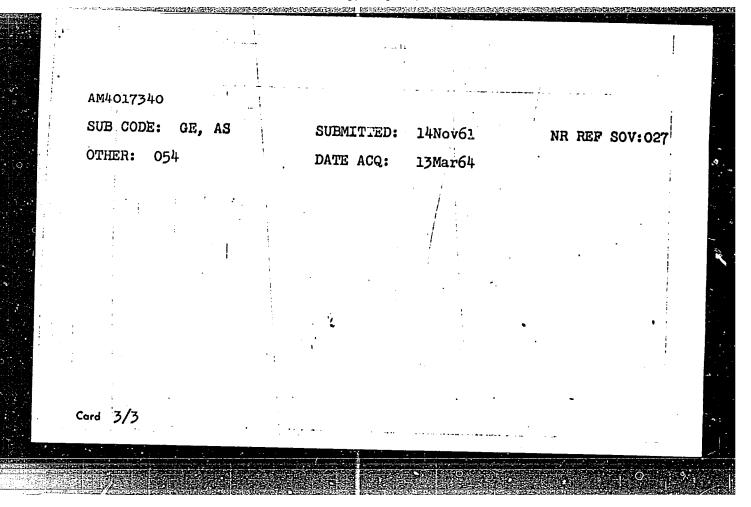


"APPROVED FOR RELEASE: Monday, July 31, 2000 CIA-RDP86-00513R000929110



THE ASSESSMENT OF THE PROPERTY OF THE PROPERTY

S/169/62/000/002/054/072 D228/D301

AUTHORS:

Kasheyev, B. L. and Lebedinets, V. N.

TITLE:

Some preliminary results of meteor activity observa-

tions in the IGY-IGC period

PERIODICAL:

Referativnyy zhurnal, Geofizika, no. 2, 1962, 3-4, anstract 2G17 (Mezhdunar. geofiz. god., Inform. byul.

no. 3, 1961, 8 - 12)

TEXT: Measurements of the number of meteors in the period from December 1957 to January 1960 are reviewed. It is mentioned that the investigation of dense central coagulations in meteor streams is significant for cosmonauts. Comparison of the corresponding pe ricds of 1958-1959 shows the diurnal variation's good repetition. The complex structure of the Geminide and Quadrantide flows was noted in the meteor-stream observations. The high activity of the Lyride flow is noted for the first time since 1922. The velocity of the Leonides was measured. / Abstracter's note: Complete translation. 7

Card 1/1

LEBELINGTE V. N.

6,4738

28702

S/021/61/000/003/006/013 D274/D301

AUTHORS:

Dudnyk, B.S., Kashcheyev, B.L. and Lebedynets', V.N.

TITLE:

Errors in radar measurements of meteor velocity,

due to diffusion

**APPROVED FOR RELEASE: Monday, July 31, 2000** 

PERIODICAL:

Akademiya nauk UkrSSR. Dopovidi, no. 3, 1961, 299-

If ambipolar diffusion is taken into account, the expression for the strength of the reflected signal at the receiver input, is

$$P_{R} = \frac{P_{T}G_{T}G_{R}\lambda^{3}\alpha^{2}}{16\pi^{2}R^{3}} \left(\frac{e^{2}}{mc^{2}}\right)^{2} e^{-2\left(\frac{2\pi r_{0}}{\lambda}\right)^{2}} |I|^{2}, \qquad (1)$$

where  $P_{\mathrm{T}}$  is the strength of the transmitter,  $G_{\mathrm{T}}$  and  $G_{\mathrm{R}}$  are the directivity coefficients of the antennas, R - the distance from the meteor

Card 1/4

 $I = \int_{-\infty}^{\infty} e^{2\pi i x^2} \cdot e^{-(x_0 - x) dx};$ (2)

CIA-RDP86-00513R000929110(

28702 \$/021/61/000/003/006/013

Errors in radar measurements...

$$x = \frac{S}{\sqrt{R\lambda}}; \quad x_0 = \frac{S_0}{\sqrt{R\lambda}}; \quad \Delta = \frac{16\pi^2 D \sqrt{R}}{V \cdot \lambda \sqrt{2}};$$
 (3)

where S is the distance along the trail from the point of mirror reflection,  $S_{\rm O}$  - the coordinate of the head of the trail, V - the meteor velocity, D - the coefficient of ambipolar diffusion. Neglecting the broadening of the trail while the principal Fresnel zones are formed, one obtains the ordinary Fresnel integral

$$I \approx e^{-\frac{16\pi^2Dt}{\lambda^2}} \int_{-\infty}^{x_p} e^{2\pi i x^2} \cdot dx.$$
 (4)

The positions of the maxima and minima of the diffraction pattern, computed by formula (4), are used for calculating the velocity of meteors, T.R. Kaiser (Ref. 1: Advances Phys., 2, 495 (1953)). The authors carried out, for various values of  $\Delta$ , numerical integration by formula (2), and determined the exact position of the maxima and minima of the diffraction pattern. Comparing them with the Card 2/4

28702 S/021/61/000/003/006/013 D274/D301

Errors in radar measurements...

results obtained by using formula (4), the errors in using approximation (4) were obtained. A figure shows the errors in velocity-values related to the function  $\Delta$  for the following velocity-measurements:  $v_1$  - measured by the distance between the first and second maximum,  $v_2$  - between first and third,  $v_3$  - between first maximum and first minimum. (The error resulting from measurements by the distance between first and second minimum ( $v_4$ ) never exceeded 2%). The figure shows that for  $\Delta$  = 1, the errors of  $v_1$  and  $v_2$  are 12.9 and 25%, respectively. For  $\lambda$  = 8 m, v = 40 km/sec, R = 200 km, to  $\Delta$  = 1 corresponds an altitude of approximately 100km. As at altitudes above 95 km, a large number of meteors is found, diffusion may lead to considerable errors in velocity measurements. Normally, the diffusion coefficient is found (according to formula (4)), by the exponential drop in the amplitude of the reflected signal. The velocity can be also found by measuring the amplitude ratio at the moments of the first and second maximum, and by the relationship between the distances between the first maximum and first minimum, and first minimum and second maximum. A special

Card 3/4

M

28702 S/021/61/000/003/006/013 D274/D301

Errors in radar measurements...

study of errors due to diffusion is being carried out by the Khar' kov Polytechnical Institute (since 1957), within the framework of the IGY. A table shows the results of measurements of 10 meteors. From formula (3) it is evident that errors due to diffusion increase with altitude and velocity of meteor; thus the error in vlis 5% at 20 km/sec at an altitude of approximately 90 km; the same error, at 60 km/sec will be at approximately 97 km altitude. There are 1 figure, 1 table and 3 references: 1 Soviet-bloc and 2 non-Soviet-bloc. The references to the English-language publications read as follows: T.R. Kaiser, Advances Phys., 2, 495 (1953); J.S. Greenhow, E.L. Neufeld, Journal Atm. Terrestr. Phys., 6, 133, (1955).

ASSOCIATION:

Kharkivs'kyy politekhnichnyy instytut im. V.I.

Lenina (Khar'kov Polytechnical Institute im. V.I.

Lenin)

PRESENTED:

by Academician V.G. Bondarchuk, AS UkrSSR

SUBMITTED:

April 9, 1960

Card 4/4

 $\chi$ 

29573

\$/033/61/038/004/007/010

E133/E135

3,2440 (1041,1395)
AUTHORS: Kashchey

Kashcheyev, B.L., Lebedinets, V.N., and Lagutin, M.F.

TITLE:

Radio echo determinations of the orbits of

individual meteors

PERIODICAL: Astronomicheskiy zhurnal, vol.38, no.4, 1961, 681-691

+ 1 plate

TEXT: The results obtained from visual observations of meteors are summarised in Ref. 1 (F.L. Whipple, Astron. J., Vol.59, 201, 1954). The radio echo method of observing meteors has been in use at Jodrell Bank since 1958 (Ref. 2; J.C. Gill, J.G. Davies, in use at Jodrell Bank since 1958 (Ref. 2; J.C. Gill, J.G. Davies, Monthly Notices Roy. Astron. Soc., Vol.116, 105, 1956). One mesult has been the discovery of large numbers of faint meteors result has been the discovery of large numbers of faint meteors (7-8 mag.) with almost circular orbits inclined at a large angle to the ecliptic (Ref. 3; Meteory, Sbornik statey, IIL (Meteors, Symposium, IIL) 1959). The lifetime of these particles must be symposium, IIL) 1959). The lifetime of these particles must be small (Ref. 4; L. Kresák, Byul. Astron. in-tov Chekhoslovakii (Bulletin Astronom. Instit. Czechoslovakia) Vol.11, 1, 1960). Apparatus was installed at Khar'kov in December 1958 for the determination of individual meteor orbits. Observations have been card 1/7

29573

Radio echo determination of the ....

S/033/61/038/004/007/010 E133/E135

made since August 1959. The general layout is indicated in Fig. 1. The radio-location equipment is at 0, 6 and 6 where 0.5 = 7100 m and 0K = 4500 m ( $\frac{1}{2}$  10 m). The transmission frequency The transmission frequency is 36.' Mc/s and the duration of the impulse is 10 microsec at 500 im alses per sec. The stations at 5 and K transmit the data they receive back to 0, after amplification. The resultant traces are photographed together. An example is shown in Fig. 2 (where the sinusoidal curve gives the Doppler frequencies determining the drift of the track). The position of the radiant and of the meteor orbit is determined by Kleiber's method (Ref.7) I.A. Kleiber, Opredeleniye orbit meteornykh potokov, SPb, 1891 (Determination of the orbit of a meteor stream)) and is done by an electronic computer; otherwise it would be impossible to reduce all the data. In order to check the accuracy of the calculated orbits, observations were made of 298 members of the Geminid stream during December 9-14, 1959. The authors first consider the braking effect of the Earth's atmosphere so that they can deduce the velocity outside the atmosphere from the observed velocity. They arrive at the equation:

Card 2/7

29573
Radio echo determination of the ... S/033/61/038/004/007/010
E133/E135

$$\Delta v_{\rm m} = -\frac{1.22}{v_{\rm o}\sigma} \tag{15}$$

which gives the velocity change in terms of the initial velocity and the parameter o (the coefficient of heat transfer). assume that  $\log \sigma = -11.2$  and that it does not vary much with the mass (Ref. 11: L.G. Jacchia, Smith, Contrib. Astroph., Vol. 2, No.9, 1958). They find that a correction of  $\sim$ 0.6 km/sec should be made, therefore, to the observed Geminid velocity. resultant r.m.s. error in the velocity measurements is \$1.8 km/sec for a single meteor. This is due to four causes; a) inaccuracies ir the allowance for atmospheric braking; b) the effect of atmospheric turbulience on velocity measurements; c) errors in velocity measurements due to diffusion of the meteor track; d) inaccuracy in the readings of the number of impulses per Fresnel zone. The data on the Geminids indicate a systematic change in the position of the radiant, and the orbital elements, with solar longitude. The authors compare their results with eptical measurements for meteors of magnitude =5 to 0 (F.I., Whipple, Ref. 1) and for meteors of magnitude 0 to +3 (Ref. 14: G.S. Hawkins, Card 3/7

Radio echo determination of the ... 29573 \$/033/61/038/004/007/010 £133/£135

R.B. Southworth, Harv. Reprint Series II-128, 1958). The average orbital elements of fifteen meteors in Ref. 1 agree with the present measures, as does the systematic change in the orbital elements. The results in Ref. 14 appear to be less accurate, but also agree with the limits of error. That there was a change in the position of the radiant was already known, but this change in the orbital elements is new. Since it appears to be connected with the mass of the particles, it can only be explained by some form of braking of the meteors (e.g. by the Poynting-Robertson effect) Previous observations (Ref, 17: B.L. Kashcheyev, V.N. Lebedinets. Astron. zh., Vol.36, 629, 1959) indicate that on the night of December 12-13 1959, a maximum was observed for meteors in the range 2-4 mag., but on the following night (13-14) the maximum was at about zero magnitude. It can be estimated from this, on the basis of the Poynting-Robertson effect, that the age of the stream is about 30 000 years (assuming a meteor density of 4 gm/cc

There are 8 figures; 3 tables and 18 references: 10 Soviet-bloc and 8 non-Soviet-bloc. The four most recent English language references read as follows:

card 4/7/

Radio echo determinations of the ... \$\frac{5}{033}\frac{61}{038}\frac{004}{007}\frac{7010}{010}\$\$ Ref.1: F.L. Whipple, Astron. J., Vol.59, 201, 1954. Ref.2: J.C. Gill, J.G. Davies, Monthly Notices Roy. Astron. Soc., Vol.116, 105, 1956.

Ref.11: L.G. Jacchia, Smith. Contrib. Astroph. Vol.2, No.9, 1958. Ref.14: G.S. Hawkins, R.B. Southworth, Harv. Reprint Series II-128, 1958.

ASSOCIATION: Khar'kovskiy politekhnicheskiy institut imeni V.I. Lenina

(Khar'kov Polytechnical Institute imeni V.I. Lenin)

SUBMITTED: July 18, 1960

Card 5/7

L3280

3,2500

5/831/62/000/008/001/016 E032/E114

AUTHORS :

Kashcheyev, B.L., Dudnik, B.S., Lagutin, M.F.,

Lebedinets, V.N., Luk'yashko, D.N., and

Lysenko, I.A.

TITLE:

Radar observations of meteors at Khar'kov

SOURCE:

Sbornik statey, Ionosfernyye issledovaniya (meteory). no.8. V razdel programmy MGG (ionosfera). Mezhduved. geofiz. kom. AN SSSR. Moscow, Izd-vo AN SSSR, 1962,

This paper reports the results of analyses of radio echoes from meteor trails which were recorded at the Khar'kovskiy politekhnicheskiy institut imeni V.I. Lenina (Khar'kov Polytechnical Institute imeni V.I. Lenin) during July 1957 - May 1959. observations were in accordance with the IGY programme and were carried out at 73.2 Mc/sec and 36.9 Mc/sec. Special measures were taken to suppress extraneous interference. Pulse lengths of ten microseconds were employed at repetition frequencies up to 500 cps and power per pulse ~50-70 kW. The detector sensitivity was 5 x  $10^{-14}$  W. The half-power beamwidth in the final Card 1/12

CIA-RDP86-00513R000929110( APPROVED FOR RELEASE: Monday, July 31, 2000

5/831/62/000/008/001/016 Radar observations of meteors at ... E032/E114

experiments was ± 20° (vertical plane) and ± 17° (horizontal plane). The meteor velocities were measured by a diffraction method in which the velocities relative to earth were determined from signal amplitude fluctuations. Altogether 300 000 reflections from sporadic meteors were recorded and average diurnal variations in the number of meteors were obtained throughout the period. Fig. 10 shows three typical distributions (number of meteors versus mean sidereal time). The velocity distributions were also determined as functions of time and are reproduced in the paper. Finally, the mass distribution of sporadic meteors was found from the lengths of the reflected pulses. It was found that

 $_{\rm s}$   $\sim$  2.  $N = N_0 m^{s-1}$ where

Owing to the large beamwidth, weak meteor showers could not be detected against the sporadic background. Brief details are given about the following showers which were the only reliably detected showers: Quadrantids, Lyrids, Geminids, n-Aquarids and Arietids There are 16 figures. (daytime).

card 2/12

CIA-RDP86-00513R000929110( **APPROVED FOR RELEASE: Monday, July 31, 2000** 

43281

s/831/62/000/008/002/016 E032/E114

3.2440 9,9600 Dudnik, B.S., Kashcheyev, B.L., and Lebedinets, V.N.

AUTHORS:

The effect of diffusion on radar measurements of the

TITLE:

SOURCE:

lonosfernyye issledovaniya (meteory). Sbornik statey, velocity of meteors no.8. V razdel programmy MGG (ionosfera). Mezhduved. geofiz. kom. AN SSSR. Moscow, Izd-vo AN SSSR, 1962,

It is noted that in all meteor velocity determinations the expansion of the meteor trail during the time of formation of the main Fresnel zones is neglected and hence the position of the maxima and minima of the diffraction pattern from which the meteor velocities are computed are found from the usual Fresnel integral

where D is the coefficient of ambipolar diffusion and  $\lambda$ wavelength. Card 1/4

APPROVED FOR RELEASE: Monday, July 31, 2000

CIA-RDP86-00513R0009291100

The effect of diffusion on radar... 5/831/62/000/008/002/016 E032/E114

T.R. Kaiser (Advances Phys., 2, 1953, 495) is said to have arrived at the erroneous conclusion that if

$$\Delta = \frac{16 \pi^2 \text{ D } \sqrt{R}}{v_{\lambda}^{\frac{3}{2}}} \leqslant 2$$

where R is the oblique range to the meteor and  $oldsymbol{v}$  is its velocity, then the approximate expression for I given above does not introduce appreciable errors into the velocity calculation. The present authors have carried out a numerical integration of the more exact expression

$$I = \int_{\infty}^{x_0} e^{2\pi i x^2} e^{-\Delta(x - x_0)} dx \qquad (2)$$

$$x_0 = \frac{s_0}{\sqrt{R\lambda}}$$

where

$$x_0 = \frac{s_0}{\sqrt{R\lambda}}$$

is the distance along the trail measured in the direction of motion of the meteor from the point of specular reflection; Card 2/4

5/831/62/000/008/002/016 E032/E114

card 3/4

The effect of diffusion on radar ... Fig.1 shows the calculated relative errors in the above approximate velocity as a function of  $\Delta$  (curve a - velocity determined from the distance between the first and second maxima; curve b - velocity determined from the distance between the first and the third maxima; curve c - velocity determined from the distance between When  $\triangle = 1$ , the errors for a, b and c are found to be 12, 9 and 25% respectively. When the velocity is determined from the distance between the second and third maxima the error is never greater than 2%. Numerical data are reproduced for meteors observed in accordance with the lGY programme. It is noted that when  $\Delta > 1.5$ , the diffusion coefficient can no longer be determined from the tail of the reflected signal because this tail is no longer exponential. However,  $\Delta$  can be found by measuring the ratio of the amplitudes at the first and second maxima. The diffusion correction reaches about 5% at velocities of 2C km/sec and heights of about 90 km when the velocity is determined from the distance between the first and second maxima. When the velocity is 60 km/sec the 5% level

The effect of diffusion on radar ... \$/831/62/000/008/002/016 E032/E114

occurs at about 97 km. Since diffusion has the maximum effect on the position of the first maximum of the diffraction pattern, it is recommended that at heights greater than 90 km it is better to use maxima other than the first maximum. There are 3 figures and 2 tables.

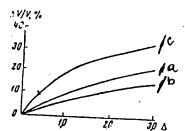


Fig. 1

Card 4/4

5/831/62/000/008/003/016 E032/E114

A radar study of the structure ...

the Jodrell Bank data (T.R. Kaiser, Meteors, London, 1950) showed that it was possible to detect masses from 0.0003 g upwards for the Geminids and from 0.0002 g upwards for the Quadrantids. only assumption for these estimates was that the meteor bodies responsible for sporadic meteors have a mass distribution of the  $N(m) = N_0(m)^{-s}$ 

with the same value of s. A preliminary report on the Geminids and Quadrantids was given in previous papers by B.L. Kashcheyev and V.N. Lebedinets (Astr. zh. 36, 1957, 623, and Astr. zh. 37, 1960, 119). The present report gives the results of new analyses in which the sporadic meteor background was taken into account more carefully. The log N versus log t curves were plotted for different times and each of the above three meteor showers (N-number of meteors with reflection lengths greater than  $\tau$ ). The data show that the central regions of meteor showers have a very complicated structure. The explanation of this structure is said to be very important to the study of disintegration of cometary nuclei and the evaluation of meteoric swarms. It was

A radar study of the structure ... \$/831/62/000/008/003/016 E032/E114

found that the Lyrids have a very narrow central region of high activity. This suggests that many showers which are considered to be of low activity may in fact contain dense particle swarms of various dimensions which may be missed in experimental studies owing to the unfavourable position of the radiant while the earth passes through the swarm. For example, it is noted that the systematic studies at Jodrell Bank did not lead to the discovery antennas employed had a very narrow directional pattern. In visual observations the probability that this short-period There are 6 figures.

Card 3/3

L 13594-63 EWG(k)/EWT(1)/FCC(w)/BDS/EEC-2/ES(v) AFFTC/ASD/ESD-3/
SSD/AFGG Pz-4/Pi-4/Pi-4 GW B/0033/63/C40/004/0719/0732
ACCESSION HR: AP3004329

AUTHOR: Lebedinets, V. N.

TITIE: Density of meteoric matter in the vicinity of earth orbit, from redio observations of meteors

BOURCE: Astronomicheskiy zhurnal, v. 40, no. 4, 1963, 719-732

TOPIC TAGS: meteor, meteor trail, meteoric particles, meteoric dust, radio-wave scatter, radar probe, meteoritic fallout, radar observation

ABSTRACT: Physical properties of meteor trails are discussed on the bests of theoretical calculations and some carlier data from radar and visual observations. Questions discussed include unique features of faint meteors, calculations on meteor detectability, and estimates of the amount of meteoritic matter falling on the earth bility, and estimates of the amount of meteoritic matter falling on the earth annually. In the discussion, an arbitrary meteoric particle is chosen for reference with a mass of 5.5 x 10-4 gram, an approach radial velocity of 40 km/sec, ence with a mass of 5.5 x 10-4 gram, an approach radial velocity of 40 km/sec, once with a mass of 5.5 x 10-4 gram, an approach radial velocity of 40 km/sec, ence with a mass of 5.5 x 10-4 gram, an approach radial velocity of 40 km/sec, of 1012 e/cm. Calculations show that the trail length of faint meteors (less of 1012 e/cm. Calculations show that the trail length of faint meteors (less than +3m absolute magnitude) will not generally exceed 17 km. The altitude has a twich maximum evaporation occurs is calculated as a function of velocity for

Card 1/2

1\_13594-63 ACCESTION ER: AP3004329

particles of two hypothetical densities, 3.5 and 2.0 g/cm3; these values of hm are shown to agree closely with measured values of meteors of similar brightness based on photographic data. Deceleration of particles in the process of evaporation and ionization is discussed and shown to be negligible for particles larger than the chosen model size. Calculations show that trail redii at the altitude under study (approximately 90-100 km) fall in the range of several meters, which also agrees with existing radar data employing the two-wavelength observation methol. The question of radio-wave scatter in meteor trails and detectability in general is treated separately for saturated and nonsaturated trails; for each case the minimum detectable return signal is expressed as a function of meteor and redar parameters. By way of illustration, a table is given which correlates the absolute magnitude of the meteor, meteor particle size, and particle velocity to probability of detection when using an 8-m radar wavelength. In order to determine the quantity of meteor particles falling on earth, the average annual incidence of meteors detected in the atmosphere is considered in terms of the model particle size; this yields an estimated annual fallout on earth of 104 to 2 x 105 tons. Orig. art. has: 5 figures, 4 tables, and 40 formulas.

ASBOCIATION: Khar'kovskiy politekhnicheskikh institut im. V. I. lenina (Khar'kov

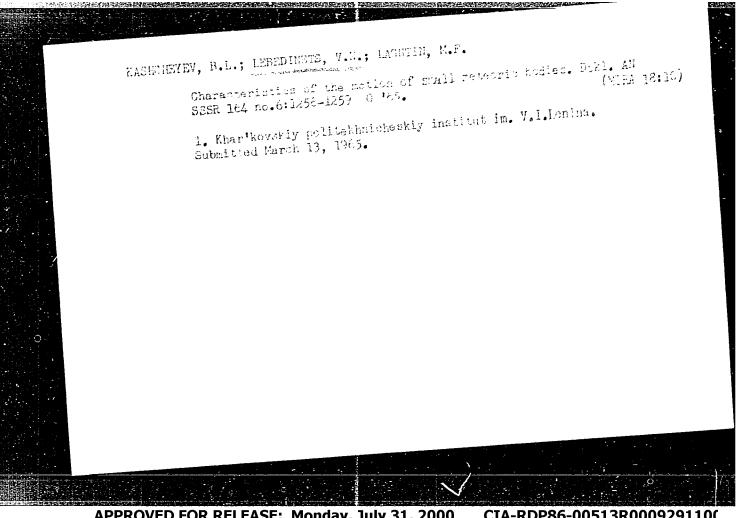
Polytechnic Institute) SUBMITTED: 06Mar62

Card 2/2.

DATE ACQ: 20Aug63 NO REF SOV: 010 SUB CODE: AS

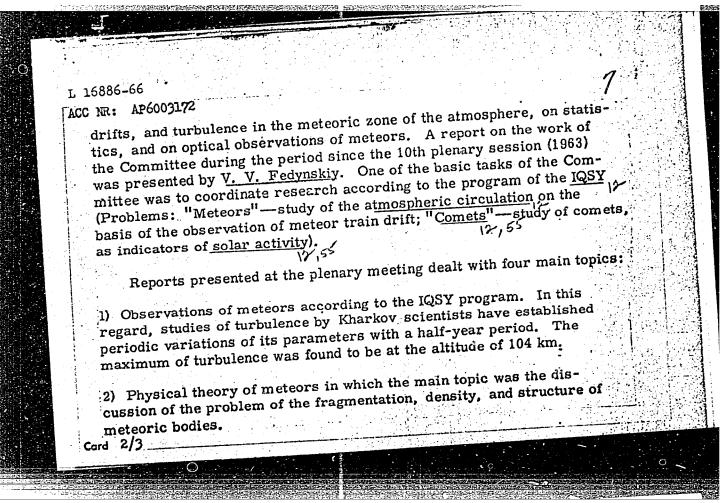
ENCL: 00 OTHER : 027

#### "APPROVED FOR RELEASE: Monday, July 31, 2000 CIA-RDP86-00513R000929110



## "APPROVED FOR RELEASE: Monday, July 31, 2000 CIA-RDP86-00513R000929110

I 16886-66 EWT(1)/FCC/EWA(d) SOURCE CODE: UR/0030/65/000/012/0103/010% AP6003172 AUTHOR: Lebodinets, V. N. (Cardidate of physico-mathematical sciences) ORG: none TITIE: At the Astronomic Council (plenary session of the Committee on Comets and Motoors) SOURCE: AN SSSR. Vestnik, no. 12, 1965, 103-104 TCPIC TAGS: astrometry, astronomic telescope, comet, meteor observation, meteor stream, meteor, astronomic conference, radar astronomy, atmospheric circulation, atmospheric turbulence ABSTRACT: The eleventh plenary session of the Committee on Comets and Meteors of the Astronomical Council, Academy of Sciences USSR, was convened at Kazan on 7-9 September 1965. Participating in it were specialists from various Soviet institutes as well as the Czechoslovak scientists, J. Raichl and J. Stol. Besides the plenary session, meetings of the sections on the physical theory of meteors and on cometary astronomy were held, as well as those of working groups on radar investigations of meteors,



L 16886-66 ACC NR: AP6003172

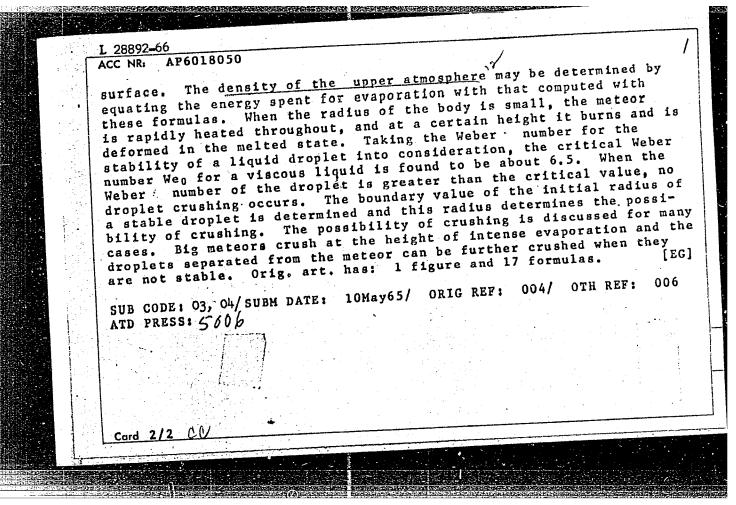
- 3) Distribution of meteoric matter in circumterrestrial space and in the solar system: analysis of the catalogue of 12,500 individual orbits of meteoric bodies, generating meteors of approximately the seventh stellar magnitude which was completed at the Kharkov Polytechnical Institute and the Institute of Applied Geophysics. About 195 meteor streams were detected, and the connection between some of them and comets was established.
- 4) Research in the field of cometary astronomy was characterized by considerable activity in observations of comets in general, but by a lag in the observation of faint comets (12-20 stellar magnitude). The Conference recommended the study of comets by the largest available telescopes, as well as the construction of special cometary telescopes aided by electronic equipment. ATD PRESS: 4177-F

SUB CODE: 03, 04 / SUBM DATE: none

Card\_3/3 B

# "APPROVED FOR RELEASE: Monday, July 31, 2000 CIA-RDP86-00513R000929110

UR/0020/66/168/003/0543/0546 EWT(1)/FCC SOURCE CODE: 3,6 AP6018050 ACC NR: 35 AUTHOR: Lebedinets, V. N.; Portnyagin, Yu. I. B Institute of Applied Geophysics (Institut prikladnoy The mechanism of crushing small meteors in the atmosphere geofiziki) TITLE AN SSSR. Doklady, v. 168, no. 3, 1966, 543-546 TOPIC TAGS: meteoric path, meteoric brightness, thermal flux, radiant, upper atmosphere ABSTRACT: Photographs revealed a series of properties of weak meteors, including path shortening, a rapid increase in brightness near the appearance point, and an anomalous braking increase in the path. These properties have been explained by the porous structure of the meteors and a hypothesis of porosity is not necessary. During the collision of a meteor with air molecules the energy is uniformly distributed on the surface of a rotating spherical meteor. The density of the thermal flux depends upon the time, the coefficient of thermal transfer, the velocity of the meteor, and the zenithal distance of the radiant. Thermal conductivity formulas were developed and their solution made it possible to determine the quantity of matter evaporated from the meteor Card 1/2



## "APPROVED FOR RELEASE: Monday, July 31, 2000 CIA-RDP86-00513R000929110

L 47481-66 EWT(1) GW  ACC NR: AP6028799 SOURCE CODE: UR/0033/66/043/004/0854/0867  AUTHOR: Lebedinets, V. N.; Kashcheyev, B. L.  ORG: Khar' kov Polytechnic Institute (Khar' kovskiy politekhnicheskiy institut); Institute of Applied Geophysics (Institut prikladnoy geofiziki)  TITLE: Meteoric matter in the vicinity of an earth orbit on the basis of radar observations of meteors	
SOURCE: Astronomicheskiy zhurnal, v. 43, no. 4, 1966, 854-866  TOPIC TAGS: radar meteor observation, meteor radiant, meteor velocity	
ABSTRACT: An annual cycle of radar measurements of individual radiants and velocities of meteors brighter than +7m was carried out at the Khar' kev Polytechnic Institute from November 1959 to December 1960. The orbits of 12,500 meteoric bodies were computed. The results obtained were corrected for the selectivity of radar observations with respect to the radiant coordinates and meteor velocities, and also by taking into account the probability of the elements of a meteoric body encountering the earth. The observed and corrected distributions of elements of the meteoric body orbits are given, and compared with the	
Card 1/2 UDC: 523, 035, 27	; 

results of photographic observations. The major axes of the orbits decrease systematically with a decrease in the mass of the meteoric bodies. The radar observations indicate two new principal types of orbits of small meteoric bodies: orbits with e < 0.7 and 30° < i < 165°, and orbits similar to those of short-period comets, differing from the latter in that they have smaller perihelion distances and dimensions. The cosmogonic significance of detecting these types of orbits is discussed. Orig. art. has: 13 figures, 8 formulas, and 2 ales.

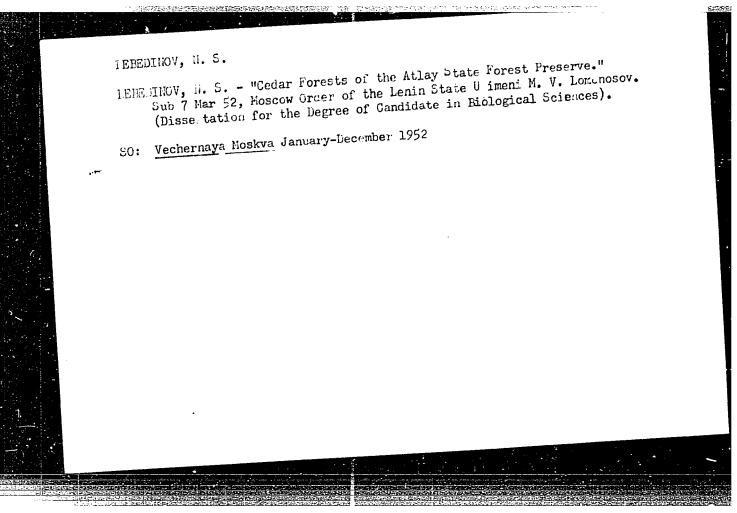
[Based on authors' abstract]

SUB CODE: 03/ SUBM DATE: 09Apr65/ ORIG REF: 006/ OTH REF: 014/

# "APPROVED FOR RELEASE: Monday, July 31, 2000 CIA-RDP86-0

### CIA-RDP86-00513R000929110

	TO THE PARTY AND AND AND ASSESSED FOR ANY AND ASSESSED FOR A STATE OF THE PARTY ASSESSED FOR A STATE OF THE PARTY ASSESSED FOR A STATE OF THE PART	
L 09	9097-67 EWT(1) GW SOURCE CODE: UR/0203/66/006/004/0712/0716	
1	UTUOR: V. N. Lebedinets and Yu. I. Portnyagin	
0	RG: Institute of Applied Geophysics (Institut prikladnoy geofiziki)	
т	TITE: Initial radius of an ionized meteor trail	
s	COURCE: Geomagnatizm i seronomiya, v. 6, no. 4, 1966, 712-716	
1	ABSTRACT: The initial radius of an ionized meteor trail has been computed, ta into account the dependence of the effective diffusion cross sections of meteo atoms in the atmosphere on the velocity of the meteor. Since the initial expandance in the atmosphere on the velocity of the first several paths of the	king rs nsion
	of the trail occurs for the most part due to the first several paths of the of the trail occurs for the most part due to the first several paths of the individual evaporating particles, the authors consider the contribution of the individual evaporating particles, the authors consider the initial radius R <sub>in</sub> satisfy to the initial radius. The derived values of the initial trail radius are the length of the free path of the initial radius of the initial radius.	paths che cora-
	radius made at Khar'kov, Kiev and Jodrell Bank. Orig. alt. Markov, Kiev and Jodrell Bank.	1d
	SUB CODE: 03 / SUBM DATE: 26Jul65 / ORIG REF: 005 / OTH REF: 006	
C	ard 1/1 nst 0925 04	038



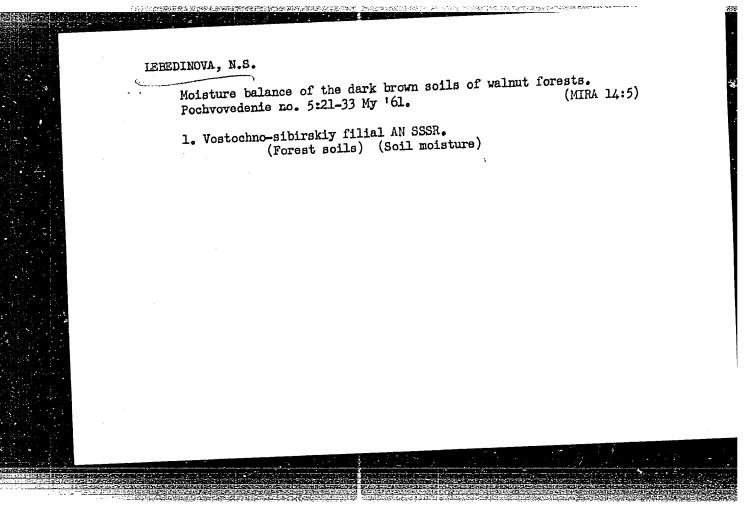
APPROVED FOR RELEASE: Monday, July 31, 2000 CIA-RDP86-00513R0009291100

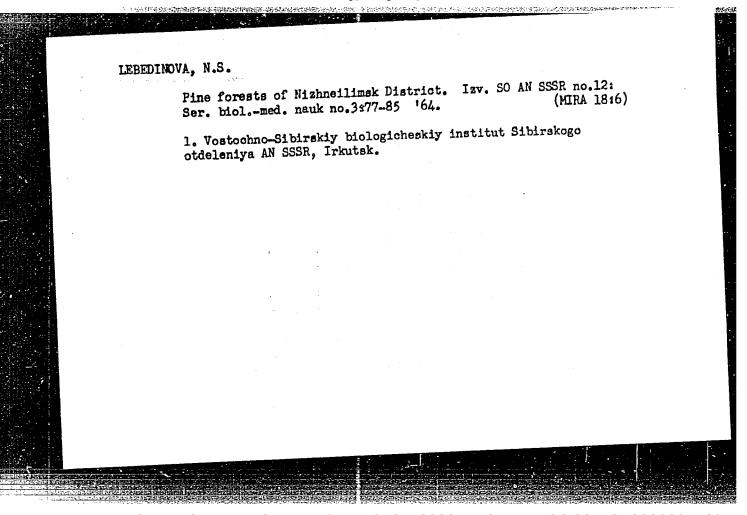
## "APPROVED FOR RELEASE: Monday, July 31, 2000 CIA-RDP86-00513R000929110

Defended his Candidates Dissertation in the Biology - Soil Faculty of Mossow State University on 7 April 1952.

Dissertation: "Gedar Forests of the Altay State National Forest."

SO: Vestnik Moskovskogo Universiteta, Seriya Fiziko-Matematicheskikh i Yestestvennykh Mauk, No. 1, Mossow, Feb 1953, pp 151-157: transl. in M-29782, 12 April 54, For off. use only.





APPROVED FOR RELEASE: Monday, July 31, 2000 CIA-RDP86-00513R0009291100

#### "APPROVED FOR RELEASE: Monday, July 31, 2000 CIA-RDP86-00513R000929110

USSR

Human and Imagel Physiology. Action of Physical Country:

Category: Factors. Ionizing Radiation.

Abs Jour: RZhBiol., No 19, 1958, 89386

Lebedinskaya, ...V. Author

On the Problem of the Effects of Fell-Out of Inst

Radioactive Isotope Strontium (Sr90). Title

Orig Pub: Med. radiologiya, 1957, 2, No 5, 22-33

The conditions of diffusion in the biosphere of Abstract:

Sr90 fall-out f llewing experimental nuclear explosions were considered and the paths of its penetration into the human organism through biclogical links were investigated; plants-animal milk -man and vegetables, fish, meat-man. Results

: 1/2 Card

T-150

Country: USSR
Category: Runn and Anical Physiology. Action of Physical
Factors. I nizin, hadiation.

Abs Jeur: RZhBiol., N. 19, 1950, 09306

of determinations of Sr90 content in human bones
in the various parts of the world were given (on
the average 0.2 strontium unit for adults and
0.5 for children). -- E.B. Glikson

Card : 2/2

### "APPROVED FOR RELEASE: Monday, July 31, 2000

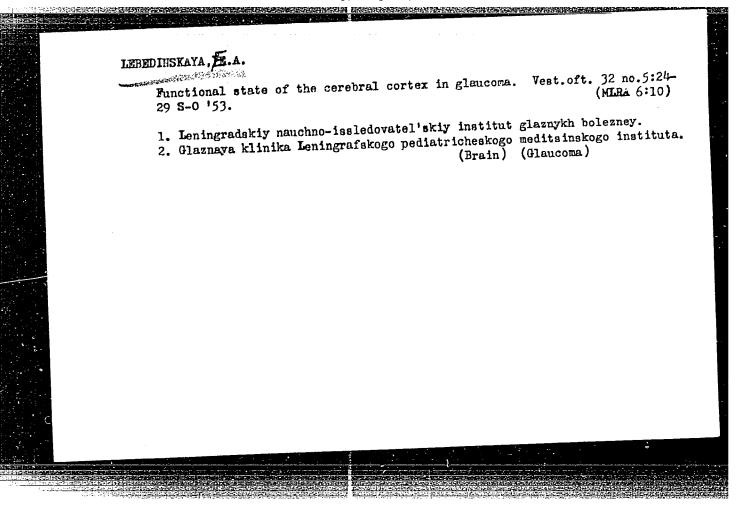
#### CIA-RDP86-00513R000929110

Country: USSR
Category: Ruman and Anical Physiology. Action of Physical Factors. I nizing Mediation.

Abs Jour: RZhBiol., No. 19, 1950, 89306

of determinations of Sr90 content in human bones in the various parts of the world were given (on the average 0.2 strontium unit for adults and 0.5 for children). -- E.B. Glikson

Card : 2/2

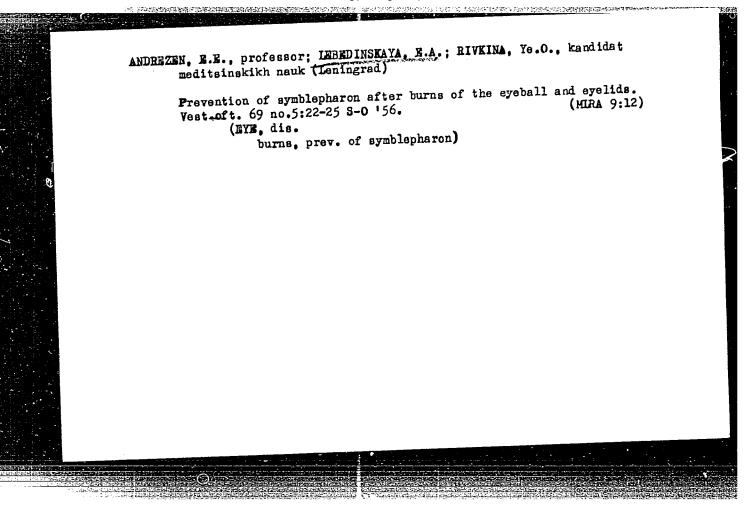


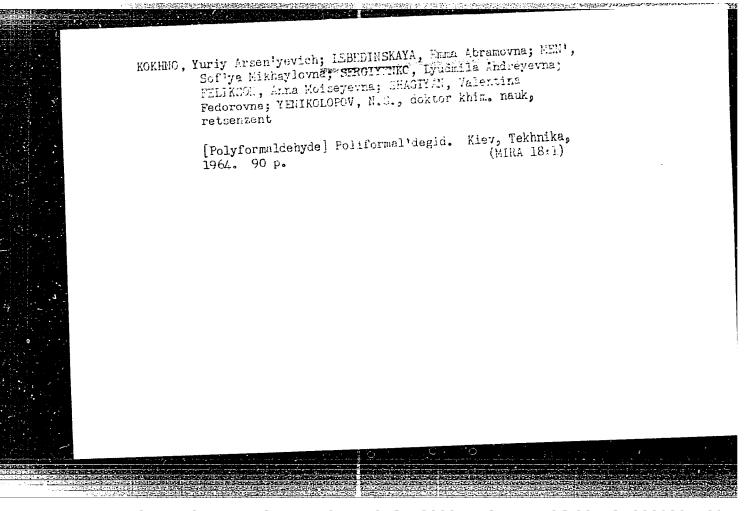
LEBEDINSKAYA, E. A.

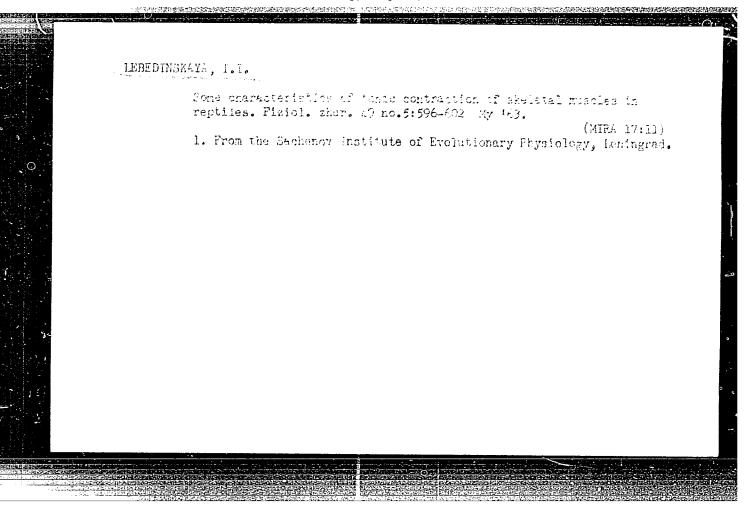
LEBEDINSKAYA, B: A.--"Materials for the functional characteristics of the Visual Analyzer as Used for Glaucoma."\*(Dissertation for Degrees in Science and Engineering Defence at USSA Higher Educational Institutions.) Min of Health Protection RSFSR, Leningrad Sanitary-Hygienic Medical Inst. Leningrad, 1955

SO: Knizhnaya Letopis! No. 25, 18 Jun 55

\* For Degree of Candidate in Medical Sciences







MIXHALEVA, O.A.; GEREDINSKAVI, I.L.

Jardian spinal reflexes from the central end of the cervical sympathetic nerve. Piziol. zhur. 51 no.1:134-141 Ja '65. (Mia 18:7)

1. Institut evolyutaiennoy fiziologii imeni Sechenova AN SSSR i Institut fiziologii imeni Pavlova AN SSSR, Leningrad.

# MIKHATEVA, O. L.; IFSGOINSKAYA, I.I.

Hechanism of the facilitation of the cardiovascular effects in the stimulation of the central caudal end of the cervical sympathetic nerve. Fiziol.zhur. 51 no.7:821-825 165.

(MIRA 18:10)

1. Institut fiziologii imeni T.P.Pavlova AN SSSR i Institut evolutsionnoy fiziologii i biokhimii imeni T.M.Sechenova AN SSSR, leningrad.

## LEBEDINSKAYA, I.I.

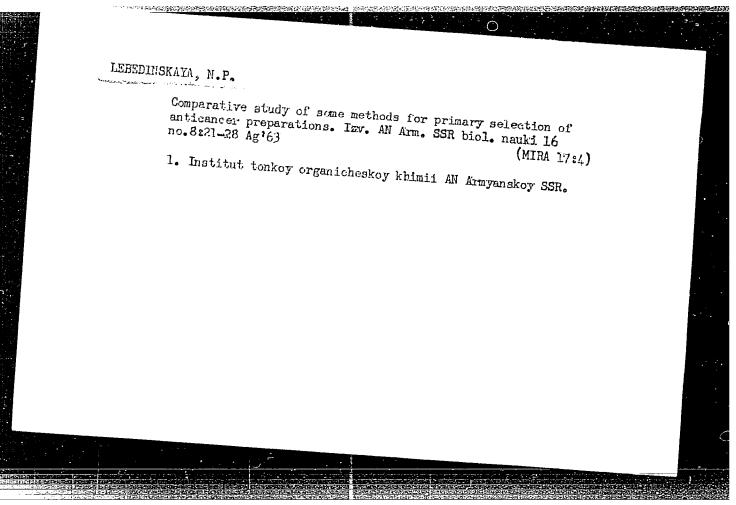
Morphological and functional differentiation in the peripheral motor apparatus in reptiles. Fiziol. zhur. 51 no.10:11991209 0 65. (MIRA 18:12)

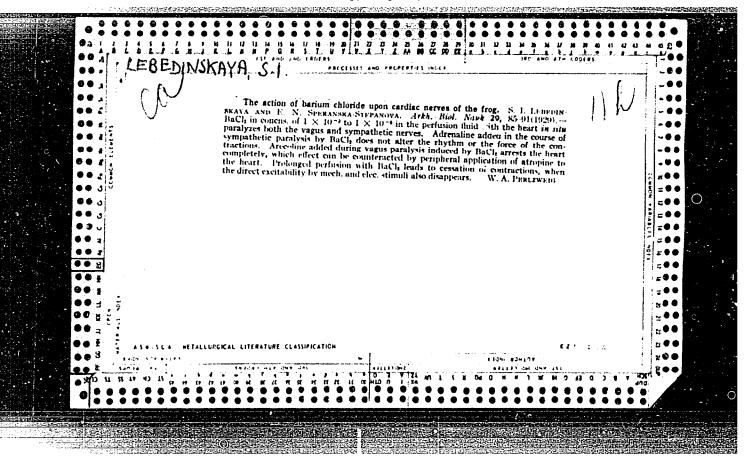
1. Institut evolutsionnoy fiziologii i biokhimii imeni I.M. Sechenova AN SSSR, Leningrad. Submitted May 11, 1964.

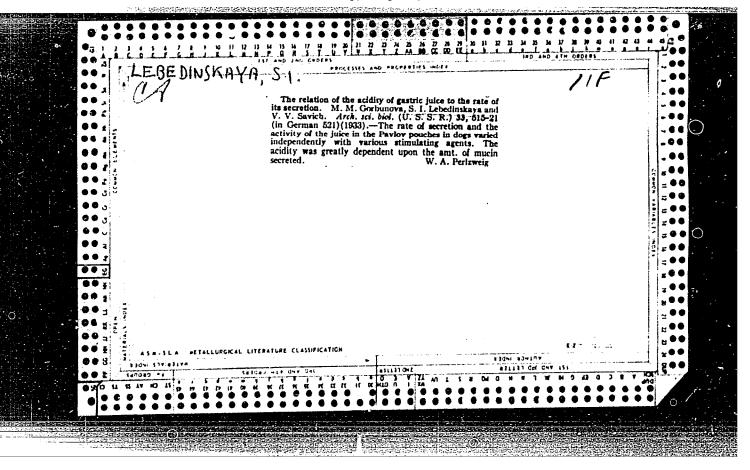
APPROVED FOR RELEASE: Monday, July 31, 2000 CIA-RDP86-00513R0009291100

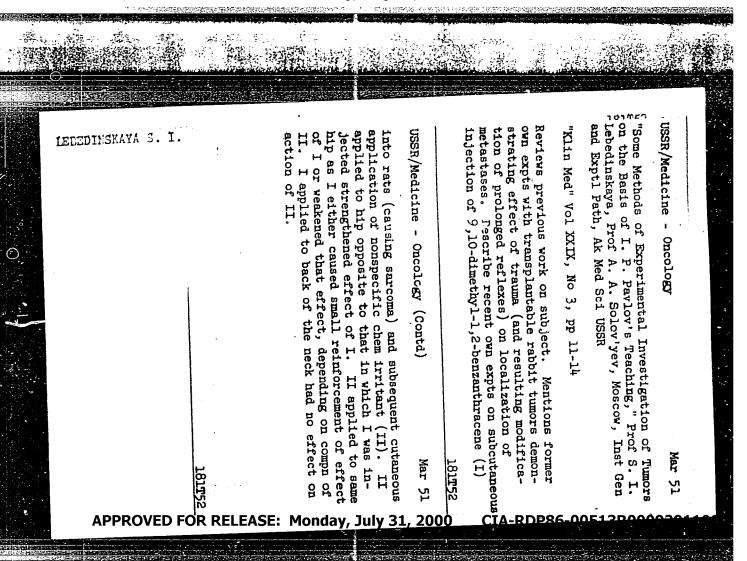
BAADE, Frits [Beade, Fritz], pref.; BATSANOVA, N.A. [translator]; FOMIN, B.S. [translator]; VISHNEV, S.M., red.; LEBEDINSKAYA, L.N., red.; KHOMYAKOV, A.D., tekhn.red.

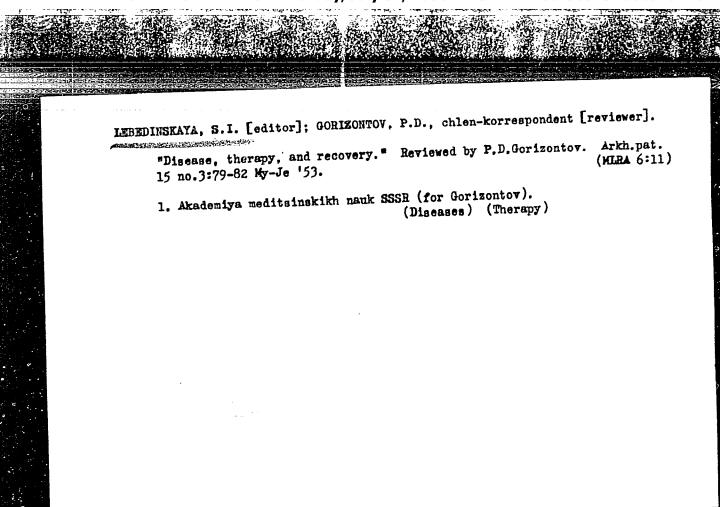
[World power engineering; nuclear power - now or in the future?]
Mirovoe energeticheskoe khoziaistvo; atomnaia energiia - seichas
ili v budushchem? Moskva, Izd-vo inostr.lit-ry, 1960. 247 p.
Translated from the German. (MIRA 13:12)

SADOMIRSKIY, D.M.; FIL'MENSHTEYN, I.D.; LEBEDINSKAYA, M.L. Gelation of latexes in the presence of polyvinyl methyl ether. Koll. zhur. 25 no.6:679-682 N-D '63. (MIRA 17:1) 1. Moskovskiy institut tonkoy khimicheskoy tekhnologii i Nauchno-issle-dovatel'skiy institut rezinovykh i lateksnykh izdeliy. 





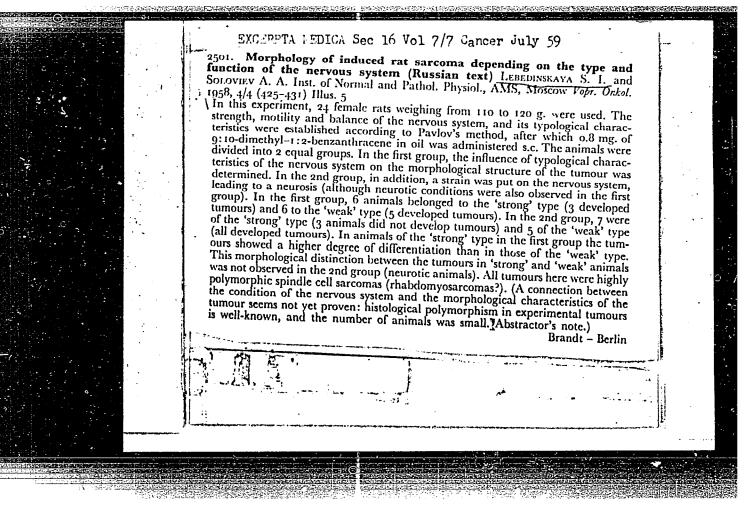




LEBEDINSKAYA, S. I.

"The Genesis and Development of Induced Tumors Under Certain Conditions of Reactivity of the Organism Altered by Supplementary Irritants" p. 322

Problema Reaktivnosti v Patologii, Medgiz, Moscow, 1954, 344pp.



The tunor processes and characteristics of its pathogenesis.
Vest.ANN SSER 14 no.7:42-50 '59. (MIRA 12:9)

1. Laboratoriya eksperimental'noy pathologii i laboratoriya eksperimental'noy patomorfologii Otdela obshchey patologii Instituta normal'noy i patologicheskoy fiziologii ANN SSER.

(MEGPLASMS etiology)

LEBEDINSKAYA, T.A., kand.med.nauk; ANANENKO, A.A., kand.med.nauk

Oxygen deficiency in pneumonias in children. Report No.1.
Pediatriia no.9:11-15 '61.

1. Iz Gosudarstvennogo nauchno-issledovatel'skogo pediatricheskogo instituta Ministerstva zdravookhraneniya RSFSR (dir. - doktor med.nauk A.F. Chernikova).

(ANOXEMIA) (PNEUMONIA)

LEBEDINSKAYA, T.A.; LEVINA, A.V.; SAVEL'YEVA, V.V.

Clinical peculiarities of staphylococcal infection originating during antibiotic treatment. Vop.okh.mat. i det. 1 no.2:61-63
Mr-Ap 156. (MIRA 9:9)

1. Iz Gosudarstvennogo nauchno-issledovatel'skogo pediatricheskogo instituta (dir.-prof. A.L.Libov, zav., klinikoy - prof.M.N.Nebytova-Luk'yanchikova) Leningrad.

(STAPHYLOCOCCUS) (ANTIBIOTICS)



Dynamics of changes in the riboflavin (vitamin B2) content of the blood of infants in intoxications of intestinal origin, pneumonia, and certain other diseases. Vop.okh.mat. i det. 1 no.2:84 Mr-Ap '56. (MLRA 9:9)

1. Iz biokhimicheskoy laboratorii i iz kliniki rannego detstva Gosudarstvennogo nauchno-issledovatel skogo pediatricheskogo instituta (dir.-prof. A.L.Libov) Leningrad.

(RIBOFLAVIN) (INFANTS-DISEASES)

in Blood #Harr Certain Hilmsses of Young Children and Its Role in the Development of Distrophy." Len, 1957, 17 pp (Len Pediatric Med Inst).

120 copies (KL, 10-58, 121).

- 39 -

KALLINIKOVA, M.B.; IESEDINSKAYA. T.A.; SHOLEBBAKOVA, M.P.

Pynamics in the change of the content of several B vitamins in the blood of small children according to various methods of administration. Pediatrila no.7.85-38 J1 '57. (MIR. 10:10)

1. Iz biokhimicheskoy laboratorii i iz kliniki rannego detstva leningradokogo nauchno-isaledovatel'skogo pediatricheskogo instituta (dir. = prof. k.L.Linov)

(VITAMINS - P)

LEBEDINSKAYA, T.A.; GROMOVA, V.N.

Cirrhosis of the liver in infants. Vop. okhr. mat. i det. 6 no. 1:88-90 Ja 161. (MIRA 14:4)

1. Iz klinicheskogo otdela (zav. - dotsent N.P. Savvatimskaya)
Nauchno-issledovatel'skogo pediatricheskogo instituta (dir.
A.P. Chernikova, zam. direktora po nauchnoy chasti - prof.
N.R. Shastin) Ministerstva zdravookhraneniya RSFSR.

(LIVER--CIRRHOSIS) (INFANTS--DISEASES)

LEBEDINSKAYA, T.A., kand.med.nauk; ANANENKO, A.A., kand.med.nauk

Some ways of compensating for oxygen deficiency in pneumonias in infants. Sow.med. 26 no.10:82-87 0 '62. (MIRA 15:12)

1. Iz kliniki rannego vozrasta i biokhimicheskoy laboratorii Gosudarstvennogo nauchno-issledovatel'skogo pediatricheskogo instituta (dir. - kand.med.nauk V.P.Spirina) Ministerstva zdravookhraneniya RSFSR, Moskva.

(PNEUMONIA) (ANOXEMIA) (VITAMIN THERAPY)

LEBEDINSKAYA, T.A., kand.med.nauk; ANANENKO, A.A., kand.med. nauk

Significance of methemoglobinemia in the pathogenesis of oxygen insufficiency in pneumonia in children. Pediatriia 4 no.7: 32-37 J1\*63 (MIRA 16:12)

1. Iz kliniki rannego vozrasta (zav. - prof. N.R.Shastin) i biokhimicheskoy laboratorii (zav. - doktor med. nauk N.Ye. Ozeretskovskaya) Nauch p-issledovatel skogo pediatricheskogo instituta (dir. - kand.med. nauk V.P.Spirina) Ministerstva zdravookhraneniya RSFSR.

LEBEDINSKAYA, T.A.; ANANENKO, A.A.

Some factors influencing the development of oxygen deficiency in pneumonia of early childhood. Cesk. pediat. 20 no.3:353-357 Mr 165

l. Gosudarstvennyy nauchno-issledovatel'skiy Pediatricheskiy institut Ministerstva Zdravookhraneniya RSFSR, g. Moskva.

FATEYEVA, Ye. M., kand. med. nauk; LEBEDINSKAYA, V. V.

"Hypotrophy of newborn infants"; clinical etiological observations. Pediatriia no.4:3-8 62. (MIRA 15:4)

1. Iz kliniki rannego vozrasta (zav. - prof. I. V. TSimbler)
Instituta pediatrii AMN SSSR (dir. - dotsent M. Ya. Studenikin)
i detskogo otdeleniya (zav. - kandidat meditsinskikh nauk Ye. Ch.
Novikova) Instituta akusherstva i ginekologii RSFSR (dir. prof. O. V. Makeyeva)

(INFANTS(NEWBORN)—GROWTH)

KALENDAROV, G.S.; LEBEDINSKAYA, Ye.I.

Apparatus for electronarcosis and method of its application in sleep therapy. Fiziol. zh. SSSR 38 no.6:751-755 Nov-Dec 1952. (CLML 23:4)

1. Laboratory of Experimental Physiology for Revival of the Organism of the Academy of Medical Sciences USSR, Moscow.

KALENDAROV, G.S.; LEBEDINSKAYA, Ye.I.

Physiological mechanism and stages of development of electronarcosis. Fiziol. zh. SSSR 39 no.2:146-152 Mar-Apr 1953. (CIML 24:3)

1. State Central Scientific-Research Institute of Physical Therapy Methods imeni I. M. Sechenov, Yalta.

LEBEDINSKAYA, Ye.I.; POLYAKOVA, A.G.

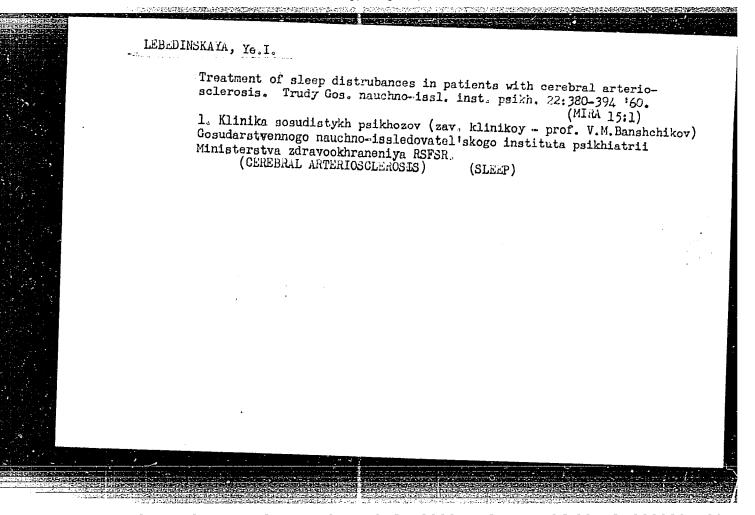
Some age-dependent changes in the interaction between the first and second signal systems in two-to seven-year-old children. Vop.psikhol. 3 no.1:53-60 Ja-F 157 (MIRA 10:3)

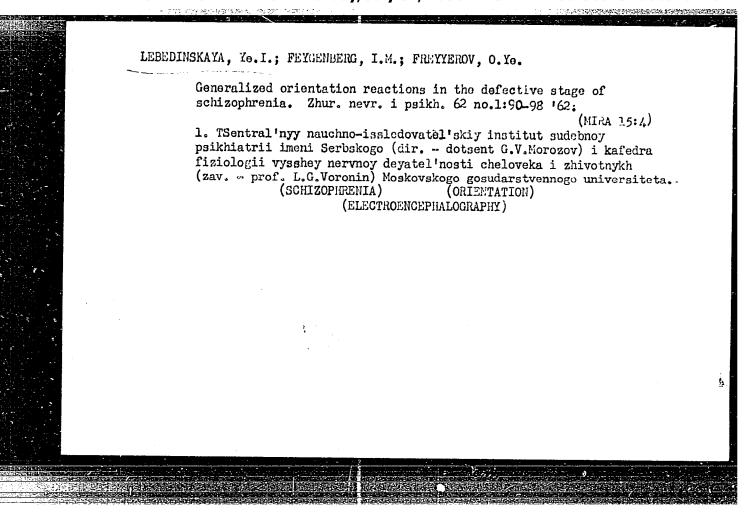
l. Kafedra fiziologii vysshey nervnoy deyatel'nosti Moskovskogo universiteta i Otdel razvitiya i vospitaniya Instituta pediatrii Akademii meditsinskikh nauk SSSR.

(Child study) (Conditioned response)

LEBEDINSKAYA, Ye. I., Cand Biol Sci-(diss) "Temporary bonds between indifferent stimuli in children from one to four years of age." Moscow, 1958. 18 pp (Mos Order of Lenin State University im M.V.Lomonosov), 170 copies. (KL, 38-58, 105).

16





Some characteristics of the formation of temperary connections in intact pigeons and birds degrived of the cerebrum according to data of electrophysiological analysis. Trudy Un. druzh. nar. 7. Vop. med. no.1:115-126

l. Kafedra normal'ncy fiziologii Universiteta Druzhby Narodov imeni Patrisa Lumumby, Moskva.

	BEDINSKAYA, Z.	of determining salt	: in canned meat.	Mias.ind.SSSR
	Hapid method (	160.	, 11 <b>(2)</b>	(MIRA 13:8)
	l. Novosibirs (Meat	kiy myasokombinat. , Canned) (	(SaltAnalysis)	
<i>t</i>				
s.				

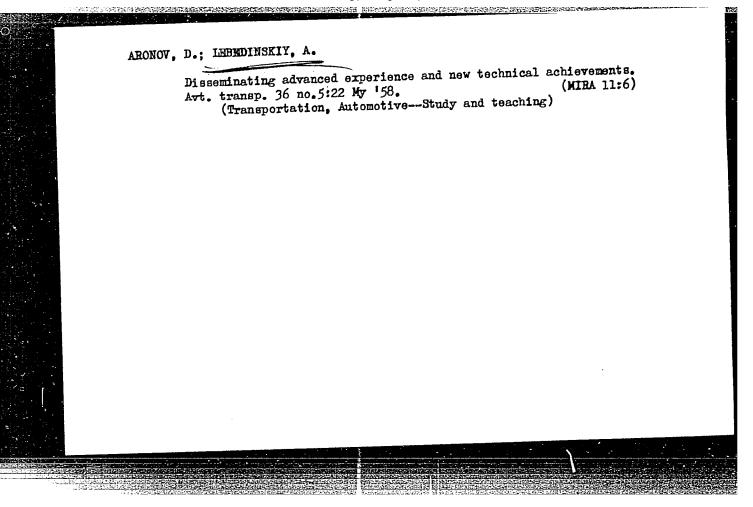
FBD/EWT(1)/EEC(k)-2/T/EWP(k)SOURCE CODE: UR/0109/66/011/004/0668/0674 AUTHOR: Korovitsyn, A. V.; Naumova, L. V.; Lebedinskaya, Z. T. ORG: none TITLE: Mode selection in the semiconcentric resonator of a gas laser 15 SOURCE: Radiotekhnika i elektronika, v. 11, no. 4, 1966, 668-674 TOPIC TAGS: gas laser, laser mode ABSTRACT: A new method of mode selection involving a special tuning of a semiconcentric resonator was investigated. The resonator was formed by spherical and plane mirrors spaced close to the spherical-mirror radius. The exact resonator length was very essential. A He-Ne laser (see figure) could, depending on the mirrors used, emit either an infrared 1.15-m or red 0.63-m radiation. It comprised: 1-discharge tube; 2-3-kv supply; Experimental He-Ne laser UDC: 621.378.335.092:621.391.84 Card 1/2

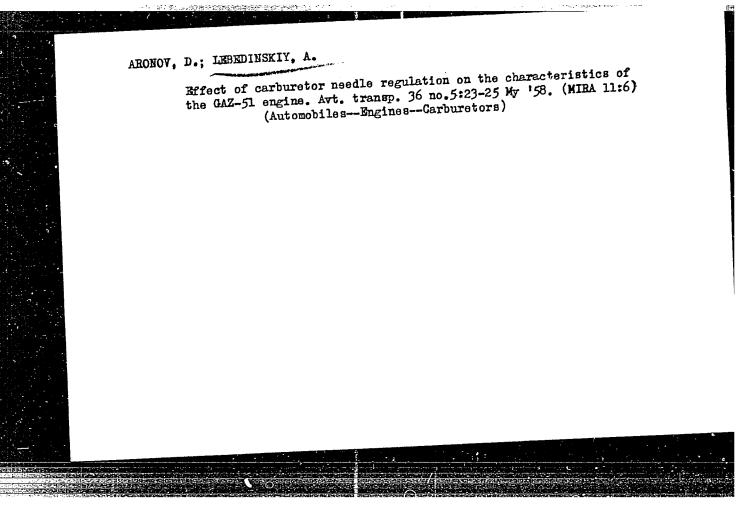
L 36194-66 ACC NR: AP6011448

3-plane mirror; 4-spherical mirror; 5-calorimeter; 6-focusing lens; 7-FEU-62 photomultiplier; 8-UR3 amplifier; 9-SCh-8 or SCh-9 spectrum analyzer. The discharge tube was 1100 mm long; the spherical-mirror radius, 1300 mm. With the inter-mirror distance considerably shorter than the spherical-mirror radius, several modes were observed. When this distance was made close to the radius, only one TEM<sub>oo</sub>-mode remained; the power of this mode constituted 0.75-0.8 maximum laser power. Thus, the mode selection is highly efficient in a near-semiconcentric system. Calculation of diffraction loss and Q-factor is presented. "The authors wish to thank A. L. Mikaelyan for guidance of this project." Orig. art. has: 8 figures and 8 formulas.

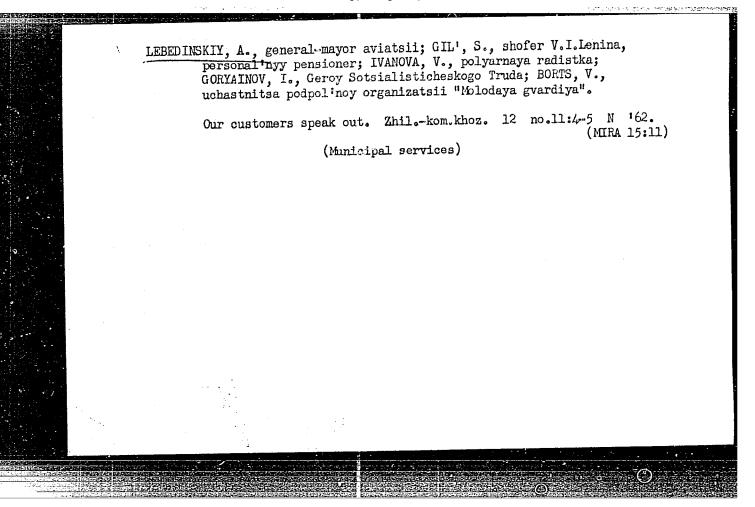
SUB CODE: 20 5 / SUBM DATE: 27Feb65 / ORIG REF: 000 / OTH REF: 004

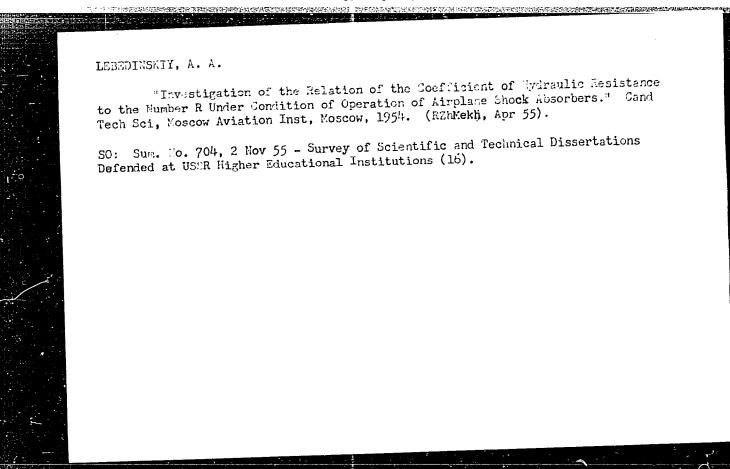
Card 2/2/1/2/





	Automatic control enters the shop. Izobr.i rats. no.9:36 S '60. (MIRA 13:10)				
\$	1. Staleprokatnyy zavod imeni Dzerzhinskogo, Odessa. (OdessaSteelworks) (Automatic control)				
4 +					





#### L 39693-65

ACCESSION NR: AP5006686

\$/0219/65/059/002/0050/0054

AUTHOR: Chernova, G. G.; Kirzon, M. V.; Safonov, V. A.; Lebedinskiy, A. B.

TITLE: The role of reflexes from the sinocarotid zone in the regulation of respiration under excessive intrapulmonary oxygen tension

SOURCE: Byulleten' eksperimental'noy biologii i meditsiny, v. 59, no. 2, 1965,

50-54

TOPIC TAGS: respiration, respiratory system, neurophysiology

ABSTRACT: The manner in which occlusion of the common carotid arteries and denervation of the sinocarotid area affect respiration in cats under excessive intrapulmonary tension (30 mm Hg) was studied. Occlusion of the common carotid arteries caused a reduction in the respiration retention occurring in response to the creation of excessive intrapulmonary oxygen tension, while denervation of the sinocarotid area caused an increase in this tension. In the case of increased pulmonary oxygen tension, occlusion of the common carotid arteries and denervation of the sinocarotid zone had no appreciable effect on the time of respiratory arrest

Card 1/2

ACCESSION NR: AP50066 86				
It was concluded that the reference on respiration under cart. has: 2 figures, 1 table	conditions of excessive intra	rea had an activating pulmonary tension. Orig		
ASSOCIATION: Kafedra fiziolo universiteta imeni M. V. Lomo State University)	ogii zhivotnykh Hoskovskogo g nosova (Department of Physio	osudarstvennogo logy of Animals, Moscow		
SUBMITTED: 16May64	ENCL: 00	SUB CODE: LS.		
NO REF SOV: 005	OTHER: 005			
를 받는 것이 되는 것들은 것을 하고 있는데 모든 생각이 되는 것을 받는데 없는데 없다. 				

LEBEDINSKY, A. F.

"A. M. Zimina & A. F. Lebedinsky: Variations of the Lechanism of Pupillary Reaction in Various Species of Animals." Received on Earch 15, 1945. (p. 305)

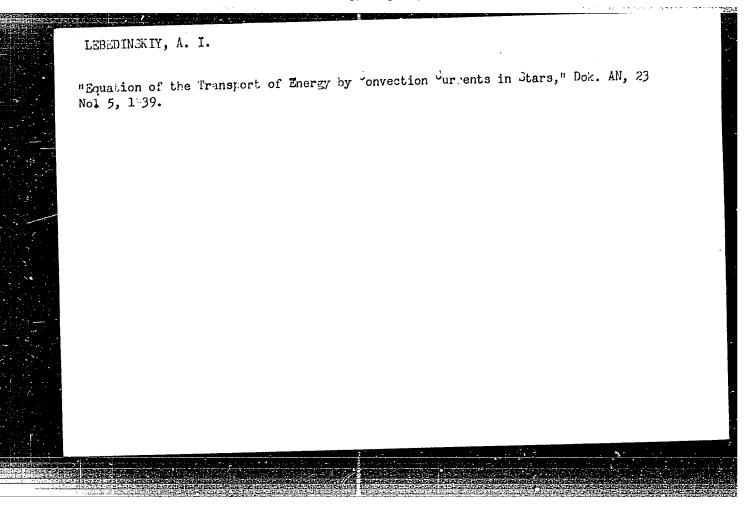
SG: Journal of General Biology, Vol. VI, contents of the issues 1-6, for 1945. No. 5

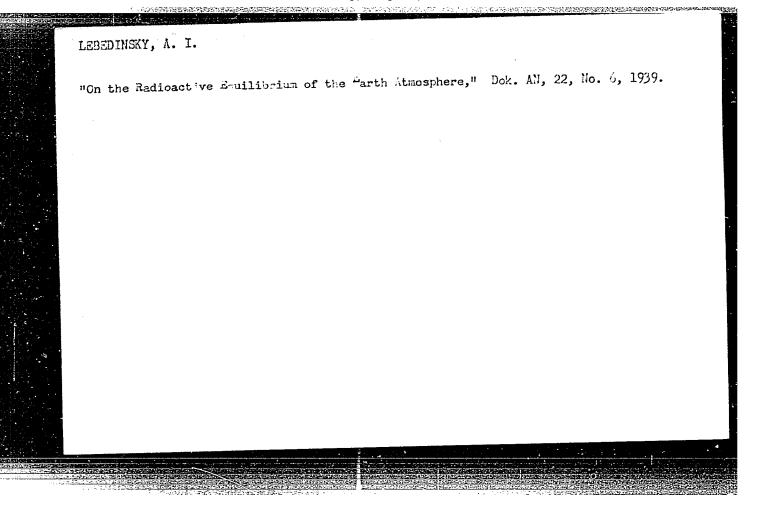
LEBEDINSKIY, A.I.; KOSHOV, S.A., prof.; SOLOV'YEVA, L.M., kand.med.nauk

On the night shift. Zdorov'e 5 no.12:21-22 D '59. (MIRA 13:4)

1. Starshiy svarshchik staleprokatnogo zavoda imeni Dzerzhinskogo,
Odessa (for Lebedinskiy).

(NICHT WORK--HYGIENIC ASPECTS)

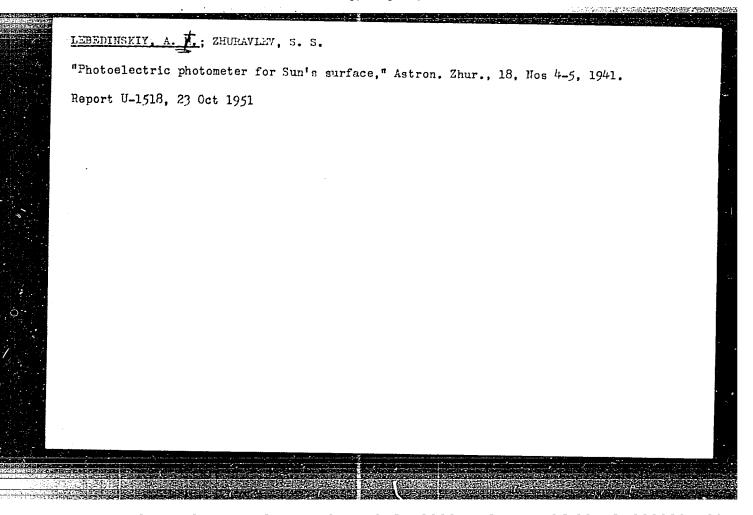




1. LEHEDINSKIY, A. I.
2. USSR (600)

"Rotation of the Sun," Astron. Zhur., 18, No 1, 1941 Astronomical observatory Leningrad State University (submitted Jun 1940)

9. Report U-1518, 23 Oct 1951



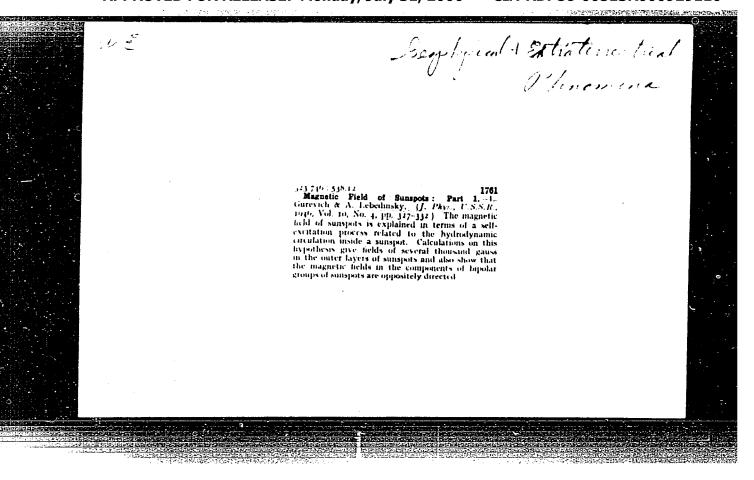
PA 4T71 LEBEDINSKY, A. I. 1945

USSR/Sunspots Magnetism, solar

"The Magnetic Field in Sunspots," L. E. Gurevich, A. I. Lebedinsky, 3 pp

"CR Acad Sci" Vol XLIX, No 2

Theory of sunspot magnetic phenomenon, which assumes that gas in sunspots flows in to the axis of symmetry of the spot in some regions, and away from it in others, which circulation of a conductive gas leads to the formation of a strong magnetic field by self-excitation from an initially weak field, i.e., the general magnetic field of the sun.



LEBEDINSKIY, A.

"Magnetic Field of Sunspots, II," Jour Phys, USSR, X, No. 5, 1946. Leningrad State U, -1946-. 6pp.

APPROVED FOR RELEASE: Monday, July 31, 2000 CIA-RDP86-00513R0009291100

LEBEDINSKIY, A. I.

"Magnetic Field of Pun Spots. I, " Zhur. Pksjer. i Teoret. Fiz., 16, No. 9, 1946. Leningrad State Univer., -1946-

### LEBEDINSKY, A.1

Lebedinsky, A. I. Structure of covelopes of novae. Astr. J. Soviet Union [Astr. Zhurnal] 23, 15-30 (1946). (English. Russian summary) [MF 16956]

The hydrodynamical equations of an extended spherical stellar atmosphere are investigated. These equations differ from the usual equations of gas dynamics in that in the equation of motion, the acceleration due to radiation pressure must be included and also the variation of temperature must be allowed for in accordance with the solution of radiative equilibrium [N. A. Kosirev, Monthly Not. Roy. Astr. Soc. 94, 430–443 (1934), equation (8); S. Chandrasekhar, ibid. 444–458 (1934), equation (51)].

Source: Mathematical Reviews,

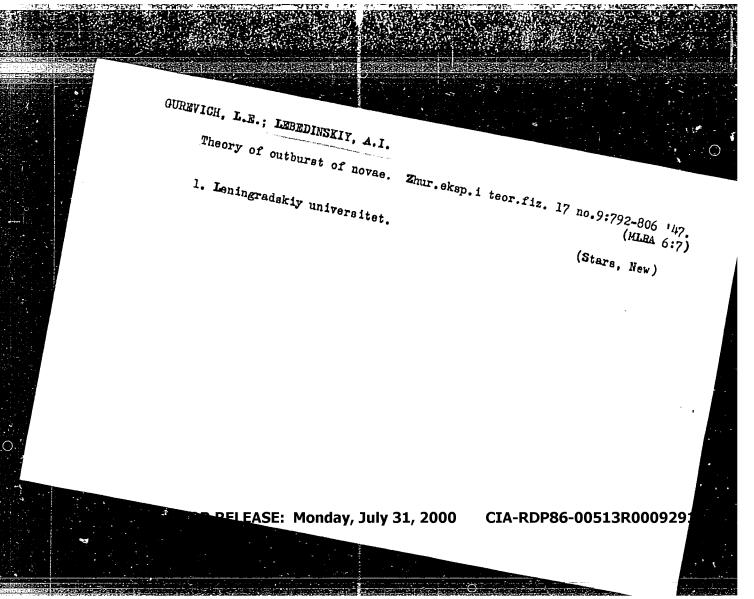
The author first shows that under certain conditions the equations admit a "characteristic surface" moving with velocity  $w=\pm (RT/\mu)^{\frac{1}{4}}$  (R the gas constant,  $\mu$  the mean molecular weight, T the temperature). He next investigates the form of the solution in the neighborhood of such a characteristic surface and shows that the structure of such a "quasi-stationary" wave can be reduced to the solution z of an equation of the form  $(z^{\frac{1}{4}}-z^{\frac{1}{4}})\partial z/\partial y+4z^{\frac{1}{4}}+z^{\frac{1}{4}}-\beta z=0$ ,  $\beta=$  constant, which tends to  $z_0$  as y=0. Here z is proportional to  $\rho T^{-\frac{1}{4}}$  and  $y=(T-T_0)/T_0$ , where  $T_0$  is the temperature on the characteristic surface. The equation is solved and the structure of envelopes of novae is interpreted in terms of the solutions.

S. Chandrasekhar.

Vol 8, No. 3

Sm god

 $\mathcal{L}^{\mathcal{L}}$ PA 16T90 LEBEDINSKIY, A. I. graphs for the purpose of studying the polarization of the solar corons. "In a Spectroscopic Laboratory," A. lebedinskiy, Yu. V. Timoreva, 2 pp to observe the eclipse 20 May 1947. At present there is under construction a set of 8 spectro-20 meters, to be ready for use by the end of 1947. In 1945 expedition went to Babayevo to obsorve the Since 1944 work carried out in completing the Spectrographic Laboratory of the Institute of Astronomy. Installation of a solar telescope factoring in the completing the spectrographic content of the co solar eclipse on 9 Jul 1945, as well as to Brazil "Vestnik Leningradskogo Universiteta" No USSR/Spectroscopic Equipment Solar phenomena 16190 Mar 1947



LEBEDINSKY, A. I.

PA ST80

USSR/Muclear Physics Stellar perturbations

Apr 1947

"Explosions in Stars, Resulting from Nuclear Reactions, as a Possible Cause for the Outbursts of Novae and Supernovae," L. Z. Gurevich, A. I. Lebedinsky, 3 pp

"CR Acad Sci" Vol LVI, No 1

An attempt to explain the outbursts of stars like  $\underline{\underline{U}}$  <u>Geminorum</u> in terms of thermal explosions brought on by nuclear reactions, as against the hypothesis of gravitational collapse.

8T8o

IEBEDINSKIY, A. I.

Apr 47

USSR/Stellar Perturbations

"Peripheral Explosions in Stars as a Result of Nuclear Reactions," L. E. Gurevich, A. I. Lebedinskiy, 4 pp

Novae and supernovae previously explained on the basis of nuclear reactions. Article gives the conditions necessary for peripheral explosions to occur. Differential and integral equations describing conditions given. Fifteen nuclear reactions (HT+H2 = He3, etc.) are given.

PA 11T80

